

PVDIS

baffle,trigger and rate

Zhiwen Zhao

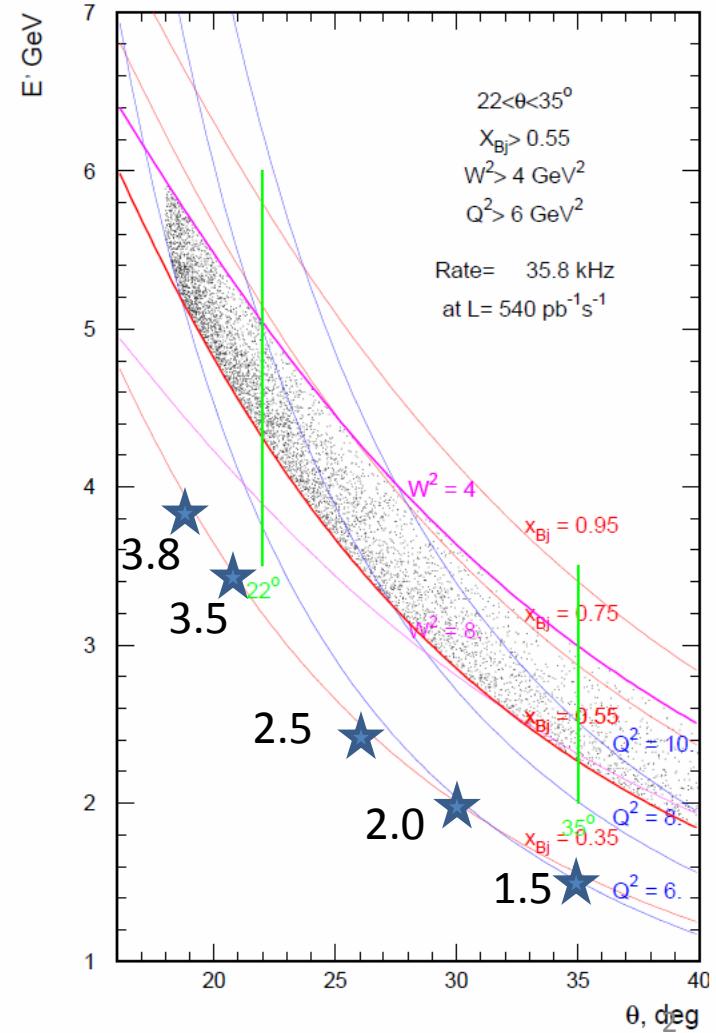
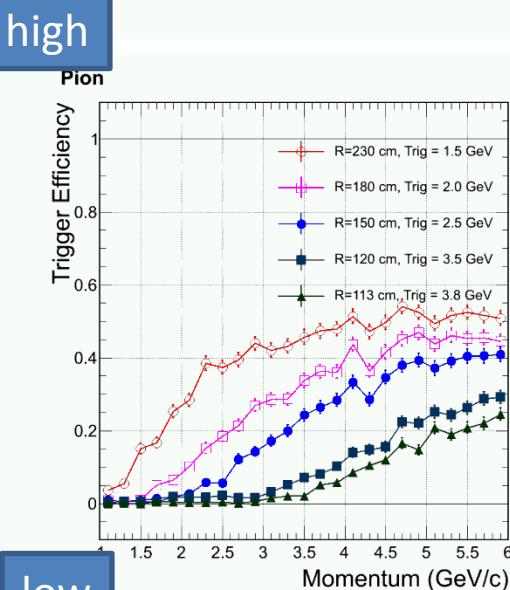
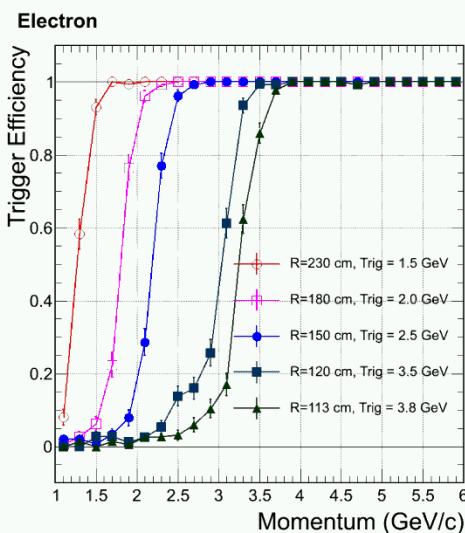
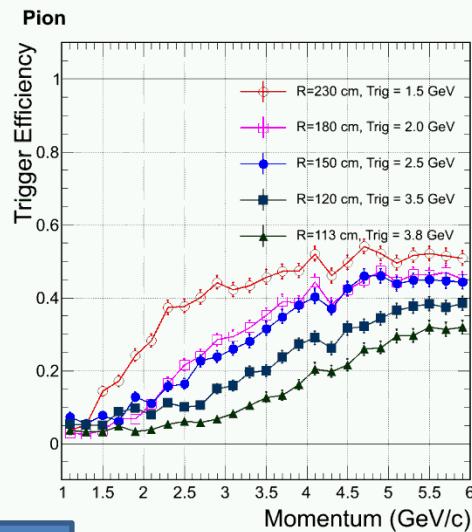
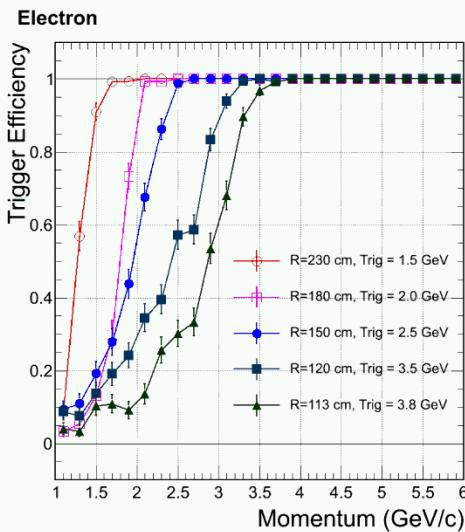
2013/09/10

2014/11/04

Assumption

More1 trig

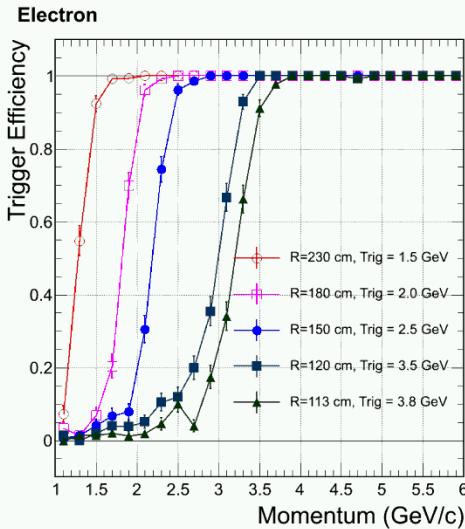
- $\text{trig_eff}(P>6\text{GeV}) = \text{trig_eff}(P=6\text{GeV})$
- $\text{trig_eff_Proton} = 0.5 * \text{trig_eff_pion}$
- $P<1\text{GeV}$, there are two methods



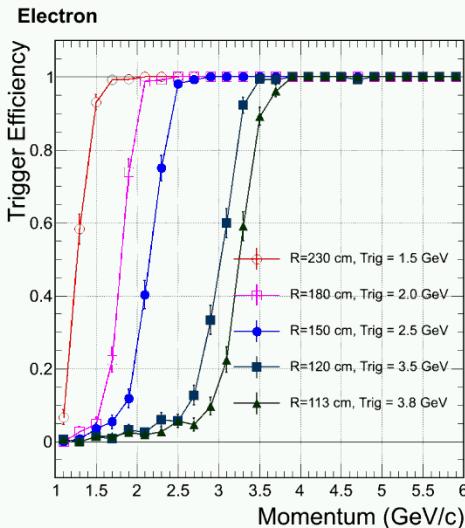
More1 block trig

Assumption

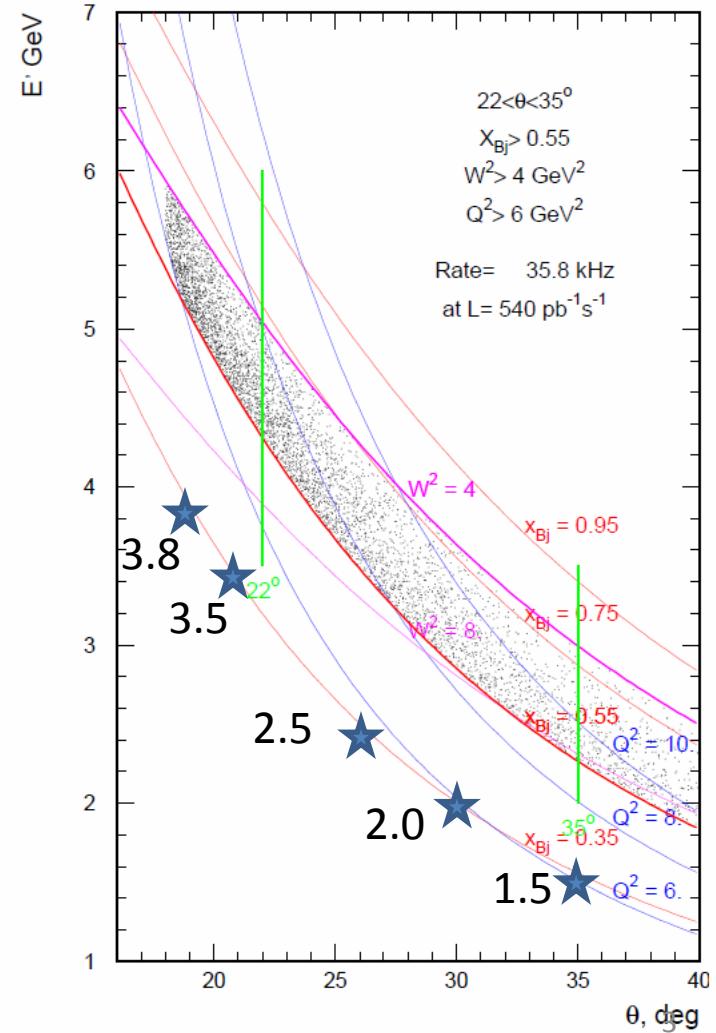
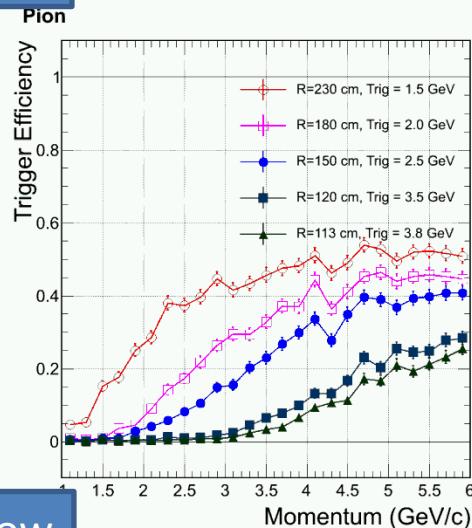
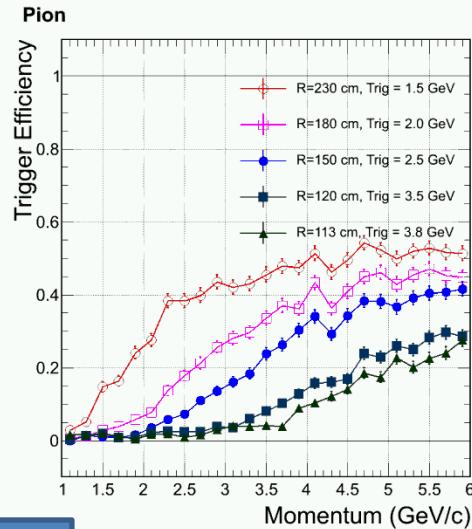
- $\text{trig_eff}(P>6\text{GeV}) = \text{trig_eff}(P=6\text{GeV})$
- $\text{trig_eff_Proton} = 0.5 * \text{trig_eff_pion}$
- $P<1\text{GeV}$, there are two methods



high



low



How to read the rate table on next slide

- All rate are in kHz
- All rate on whole EC plane, divide by 30 to get sector rate
- Top section is without trig cut, bottom section is with trig cut
- Rate has distribution over phi angle every 12 degree, we take 0-6 degree as high rate area and 6-12 degree as low rate area. The full rate area includes both

Rate (kHz)		More1 block(more1 block trig)			More1 (more1 trig)		
		full	High	Low	full	High	Low
e (W>2)	kry	413	149	265	469	205	264
π^-	Kry				5.29e4	4.11e4	1.19e4
	lead	5.08e5	2.72e5	2.36e5	5.45e5	3.15e5	2.30e5
π^+	Kry				2.27e4	1.00e4	1.27e4
	lead	2.13e5	0.98e5	1.15e5	2.13e5	1.04e5	1.08e5
$\gamma(\pi^0)$	kry	5.06e3	5.06e3	0	1.20e6	1.20e6	0
	lead	8.44e7	4.16e7	4.28e7	7.51e7	3.37e7	4.14e7
p	Kry				3.65e3	1.80e3	1.85e3
	lead	5.50e4	2.38e4	3.12e4	5.49e4	2.65e4	2.84e4
Total (lead)							
e (W>2)	kry	311	80	231	359	128	231
π^-	Kry						
	lead	5.88e3	3.80e3	2.08e3	1.37e4	1.17e4	1.93e3
π^+	Kry						
	lead	0.57e3	0.19e3	0.38e3	1.54e3	1.18e3	0.36e3
$\gamma(\pi^0)$	kry	2.1	2.1	0	1.23e4	1.23e4	0
	lead	26	26	0	1.31e4	1.31e4	0
p	Kry						
	lead	0.24e3	0.12e3	0.12e3	0.61e3	0.45e3	0.16e3
Total (lead)		7.03e3	4.22e3	2.81e3	2.93e4	2.66e4	0.27e4

Method 1

$$\text{trig_eff}(P < 0.5\text{GeV}) = 0$$

$$\text{trig_eff}(0.5 < P < 1\text{GeV}) = 0.5 * \text{trig_eff}(P = 1\text{GeV})$$

Rate (kHz)		More1 block(more1 block trig)			More1 (more1 trig)		
		full	High	Low	full	High	Low
e (W>2)	kry	413	149	265	469	205	264
π^-	Kry				5.29e4	4.11e4	1.19e4
	lead	5.08e5	2.72e5	2.36e5	5.45e5	3.15e5	2.30e5
π^+	Kry				2.27e4	1.00e4	1.27e4
	lead	2.13e5	0.98e5	1.15e5	2.13e5	1.04e5	1.08e5
$\gamma(\pi^0)$	kry	5.06e3	5.06e3	0	1.20e6	1.20e6	0
	lead	8.44e7	4.16e7	4.28e7	7.51e7	3.37e7	4.14e7
p	Kry				3.65e3	1.80e3	1.85e3
	lead	5.50e4	2.38e4	3.12e4	5.49e4	2.65e4	2.84e4
Total (lead)							
e (W>2)	kry	311	80	231	359	128	231
π^-	Kry						
	lead	4.83e3	3.43e3	1.40e3	1.08e4	9.34e3	1.41e3
π^+	Kry						
	lead	0.28e3	0.11e3	0.17e3	0.62e3	0.44e3	0.18e3
$\gamma(\pi^0)$	kry						
	lead	4	4	0	4.72e3	4.72e3	0
p	Kry						
	lead	0.18e3	0.10e3	0.08e3	0.44e3	0.32e3	0.12e3
Total (lead)		5.61e3	3.72e3	1.88e3	1.69e4	1.50e4	0.19e4

Method 2

$\text{trig_eff}(P < 1\text{GeV}) = 0$
The trig rate in this low energy region is estimated separately.

Trigger Rate Summary

- Max sector DAQ rate 65kHz, if it's 30kHz then, it cost 2.1M\$ less, 20kHz for 3.7M\$ less according to Alex's DAQ talk in May 2013 Collaboration meeting.
- Cherenkov will help to reject trigger from hadrons background (random trigger) with 30ns window, but it has no effect on true electron
- Trigger rate per sector
 - Random trigger per sector (method 2)
$$((4.83e3+0.28e3+4+0.18e3)+0.1e3*30)*2e6*30e-9/30=16.6\text{kHz}$$
 - True electron trigger per sector
$$311/30= 10.4 \text{ kHz}$$
 - Total $16.6+10.4=27\text{kHz}$

Err_Apv(%)

x	0.20-0.30	0.30-0.35	0.35-0.40	0.40-0.45	0.45-0.50	0.50-0.55	0.55-0.60	0.60-0.67	0.67-0.80
More1 Before trig cut	0.262	0.284	0.275	0.286	0.314	0.354	0.427	0.468	0.641
More1 after trig cut	0.364	0.315	0.283	0.288	0.314	0.354	0.427	0.468	0.641
More1 block before trig cut	0.290	0.304	0.287	0.294	0.319	0.356	0.427	0.468	0.641
More1 block after trig cut	0.422	0.333	0.294	0.296	0.32	0.356	0.428	0.468	0.641

New trigger strategy

- ▶ Embedding bgd stochastically according to its 3D distribution
- ▶ Look for percentage of 3ons trigger window that pass trigger threshold
- ▶ Good for low energy background pile ups
- ▶ Can not handle rare events due to stat. limit
- ▶ Handle $p < 1\text{GeV}$ background particle trigger
- ▶ Embedding bgd stochastically according to its 3D distribution
- ▶ Produce trigger turn on curve for high energy particle
- ▶ Good for rare events, e.g. DIS
- ▶ Can not handle low energy particle dominated trigger, which is non-linear
- ▶ Handle $p > 1\text{GeV}$ particle dominated trigger

From background embedding

From trigger turn on curve

Low energy ($P < 1\text{GeV}$) particle trigger

- ▶ Place a calorimeter 6+1 cluster at given reference radius location
- ▶ Assume a 30ns trigger integration window, stochastically simulate which bgd particle would fly into calorimeter
 - including e/gamma/pi+/-/o/proton, 1keV – 1GeV
 - Particle with $P > 1\text{GeV}$ is ignored in this case, since their trigger rate should be counted in high energy trigger curve x rate study
- ▶ Simulate scintillator energy deposition in the shower part for all these particle and sum to give a trigger signal
- ▶ Repeat for 60k times, check the probability to produce a trigger. Trigger threshold set according to the radius
- ▶ Multiply by number of trigger channels and get the total low energy trigger rate

Inspect on few triggered case

- ▶ For low radiation slice at R=230 cm, trigger threshold is
 - scintillator energy > 283 MeV
 - targeted high trigger efficiency for electron with E>1.5 GeV
- ▶ 9 out of 60k simulations produced a trigger

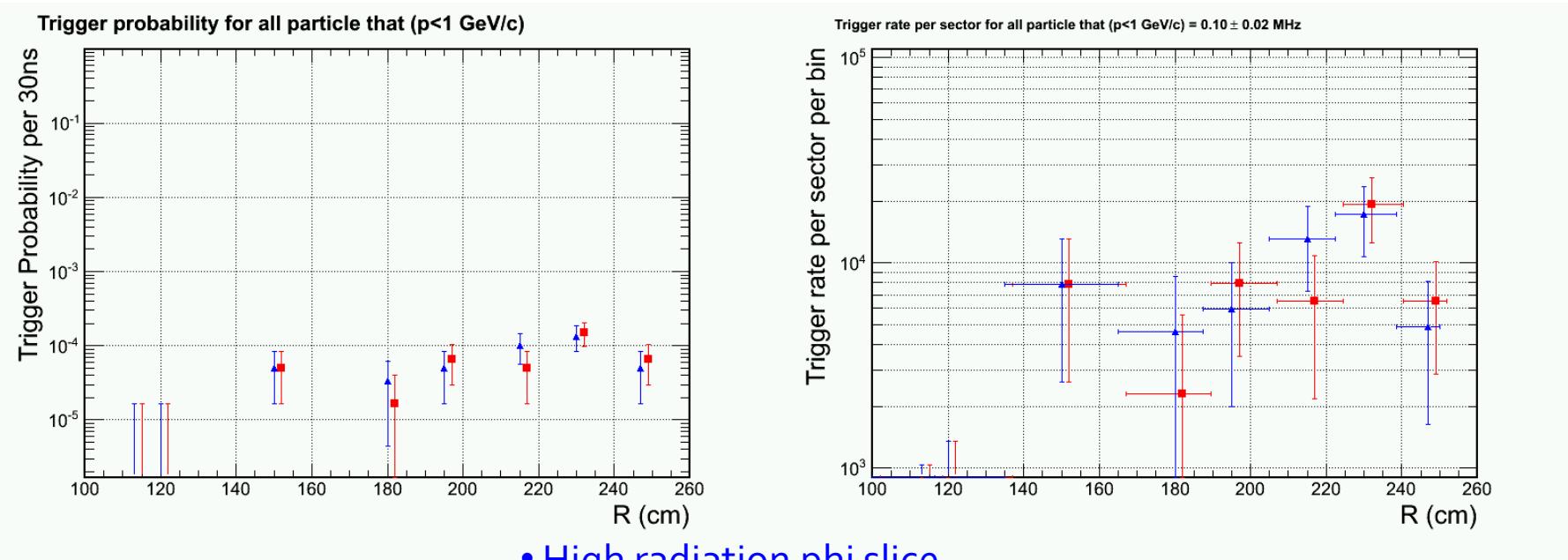
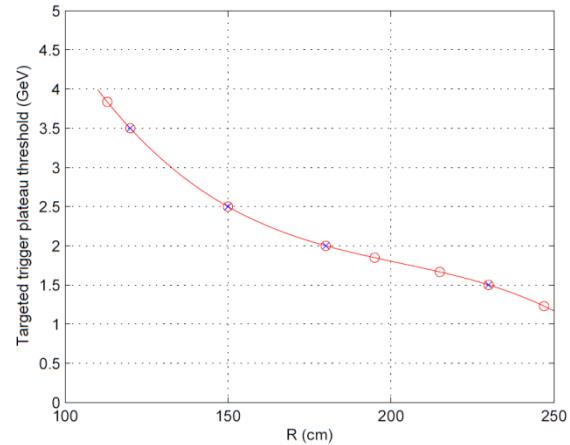
n_* : number of that particle for 30ns window

sh_*: shower scintillator energy deposition for that particle species

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*****  
*   Row   *   n_elec *   sh_elec *   n_gamma *   sh_gamma *   n_gamma_p *   sh_gamma_*   n_pip *   sh_pip *   n_pim *   sh_pim *  
*****  
*   5116   *   0   *   0   *   11   *   1.4133610   *   4   *   0.8096210   *   0   *   0   *   1   *   287.67065 <- Pi- dominated  
*   10508   *   0   *   0   *   13   *   0   *   10   *   17.858110   *   0   *   0   *   1   *   272.90136 *  
*   12082   *   0   *   0   *   13   *   1.1497589   *   5   *   3.1542911   *   1   *   328.03814   *   0   *   0 <- Pi+ dominated  
*   26961   *   0   *   0   *   15   *   0   *   9   *   13.370458   *   0   *   0   *   1   *   277.56695 *  
*   31170   *   0   *   0   *   18   *   3.771492   *   4   *   3.6389594   *   0   *   0   *   1   *   301.99948 *  
*   37962   *   0   *   0   *   12   *   0   *   2   *   0   *   0   *   0   *   2   *   315.43313 *  
*   40813   *   0   *   0   *   20   *   10.953822   *   6   *   12.016947   *   0   *   0   *   1   *   266.20440 *  
*   42284   *   1   *   0   *   13   *   1.1786102   *   5   *   1.1385887   *   1   *   82.557189   *   1   *   216.75323 <- two pion pile up  
*   42872   *   0   *   0   *   16   *   0.9754827   *   4   *   0   *   0   *   0   *   1   *   285.33731 *  
*****
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Total rate from P<1GeV particle

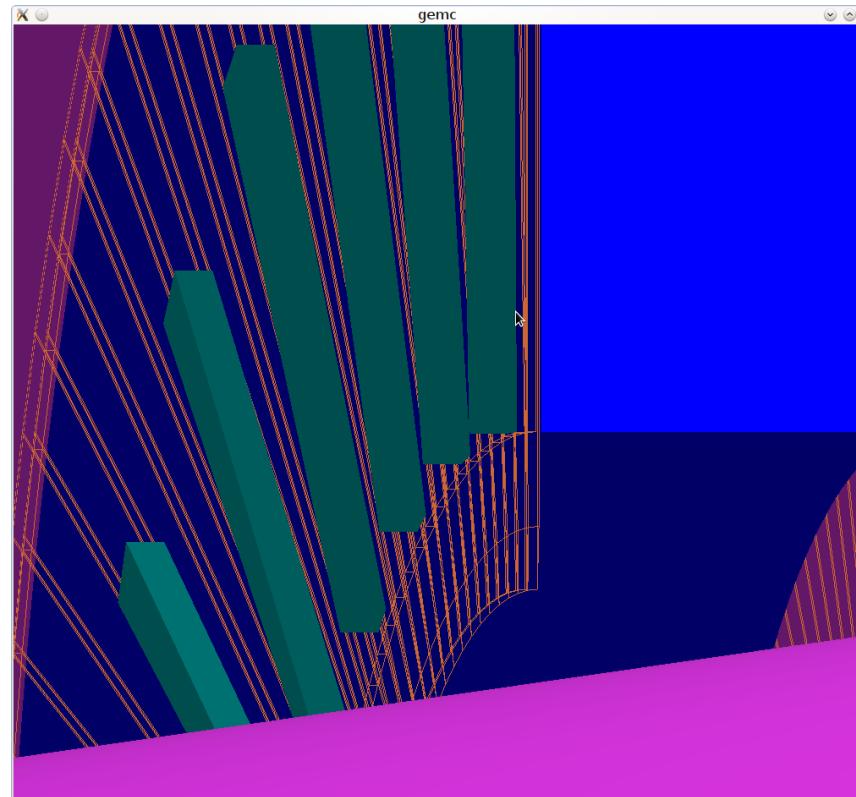
- ▶ Sum = 0.10 ± 0.02 MHz per sector
 - Statistical precision can be improved with more simulation
 - Ignored correlation between neighboring trigger channels -> over estimate
- ▶ Dominated by radius region $R \sim 230$ cm, where trigger threshold is low ($E_{\text{target}} = 1.5$ GeV)



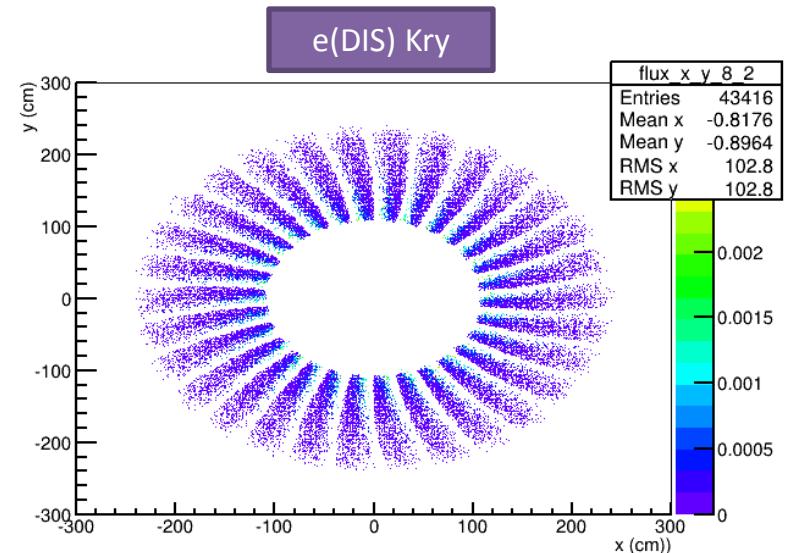
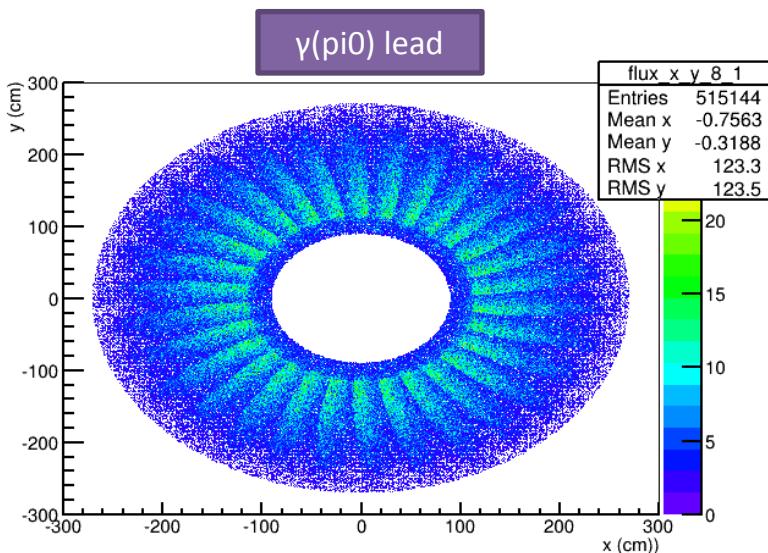
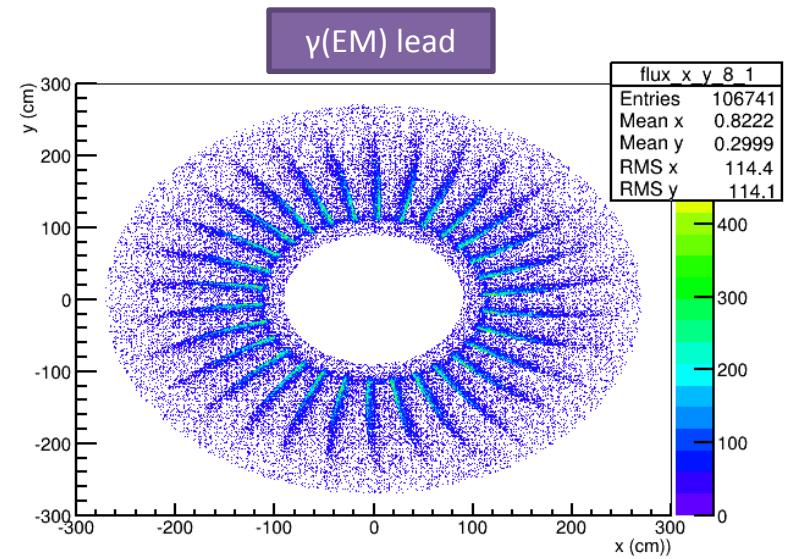
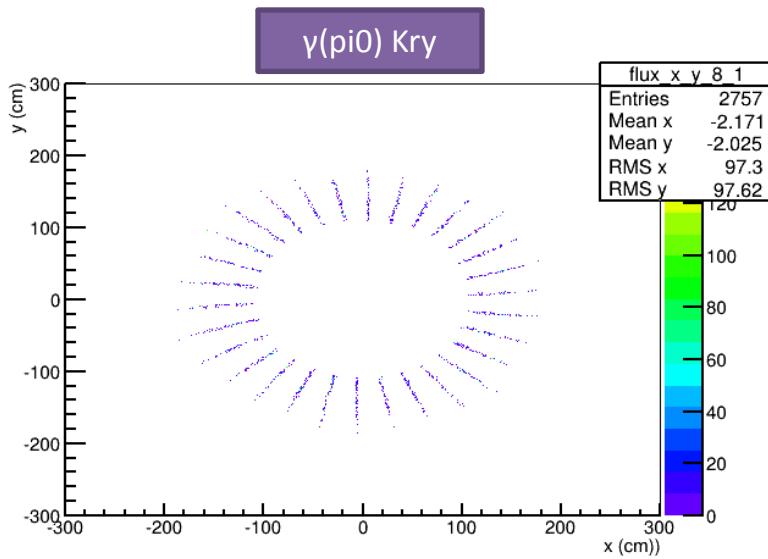
- High radiation phi slice
- Low radiation phi slice

EC photon block

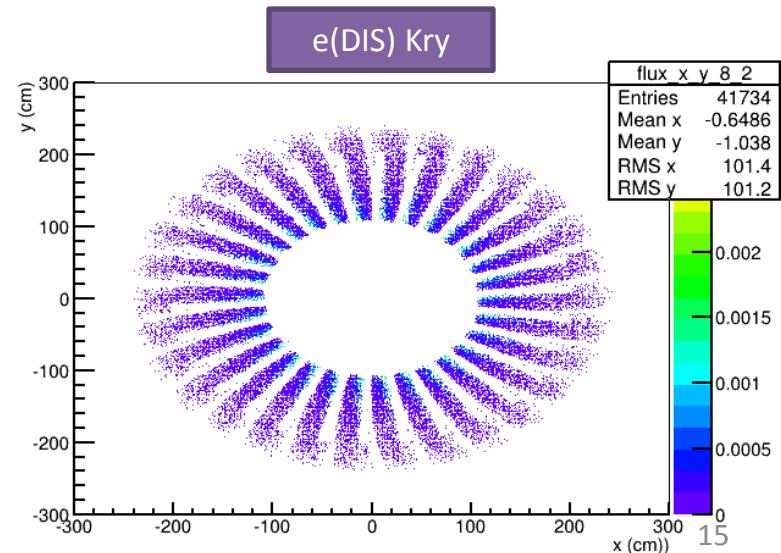
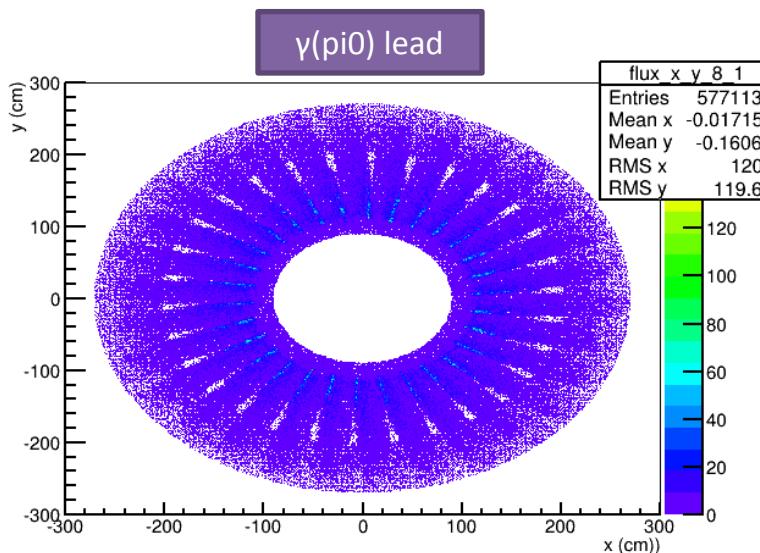
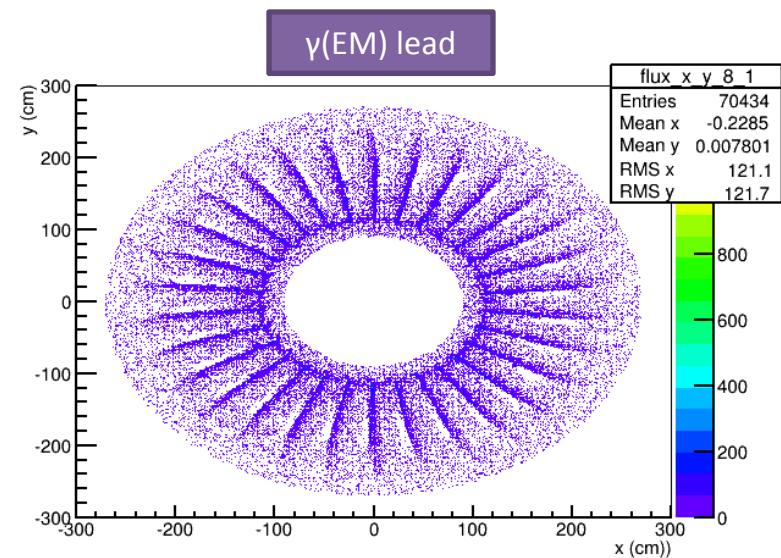
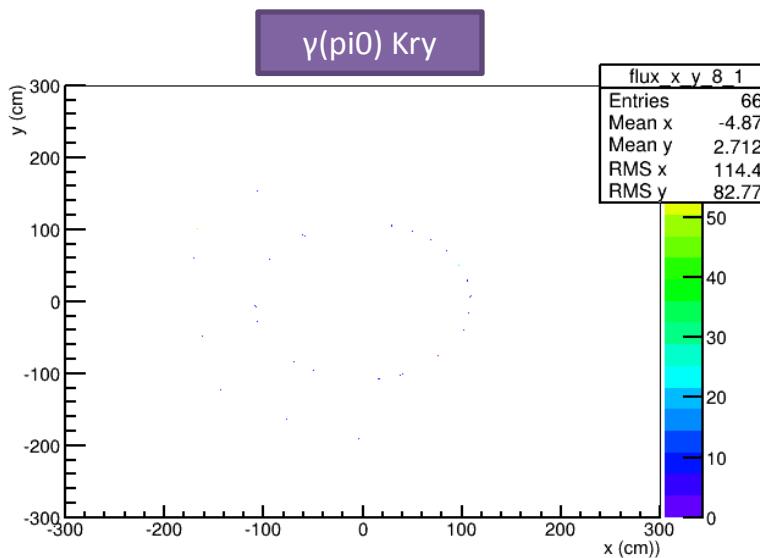
- EC module R(110,265)cm
- EC photon block (“more1 block”)
 - 30 of them
 - R(110-200)cm
 - Start from 2.2 degree and width 2.5 degree. (They can be further optimized)
 - 5cm($8*X_0$) thick lead, hope to reduce photon energy by 1 order
- EC module reduction about 180, 10% of total



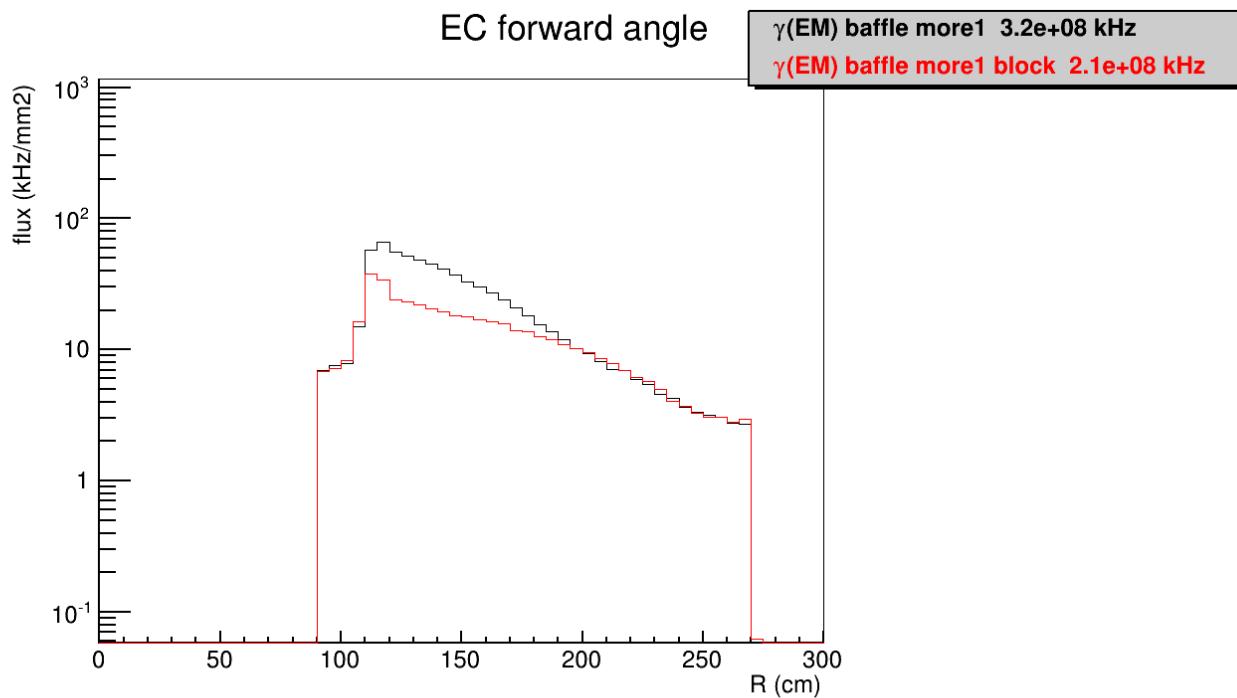
Rate of More1



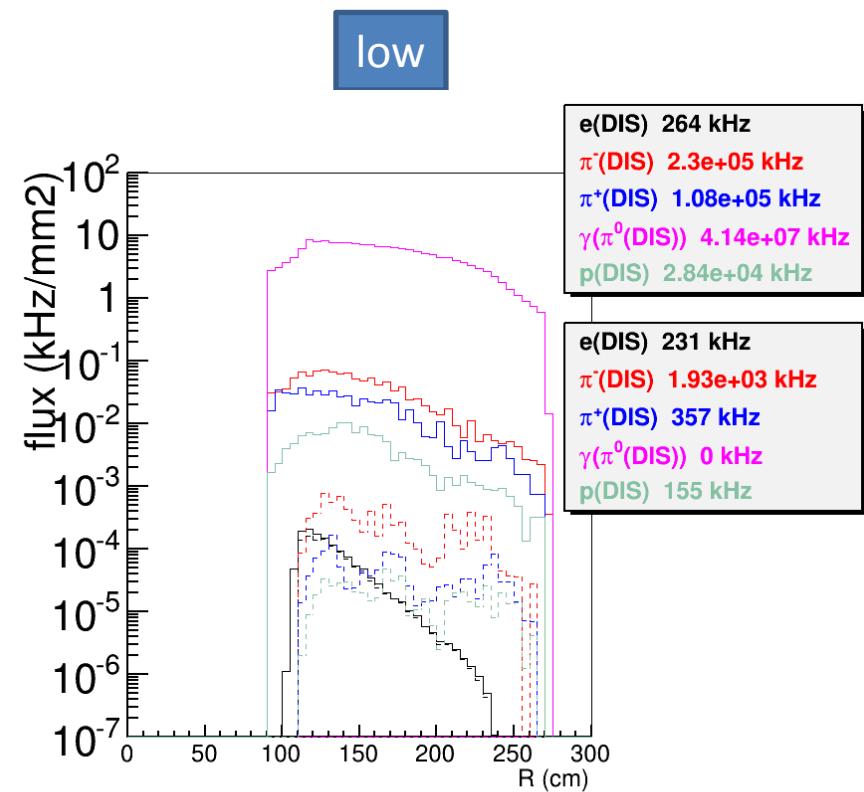
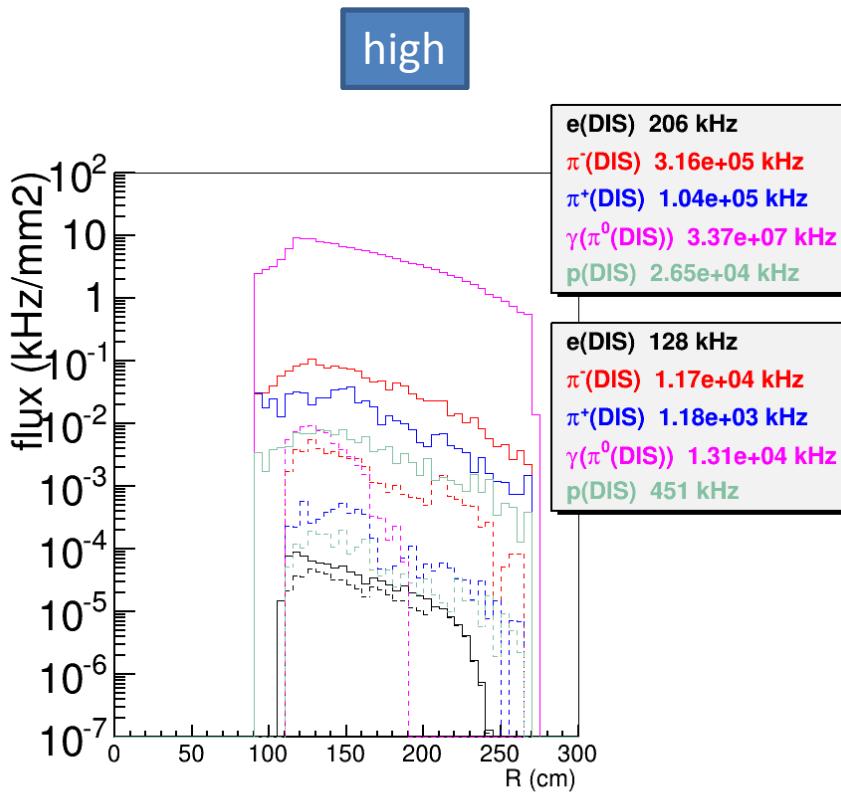
Rate of More1 block



γ (EM) lead rate of more1 and more1 block

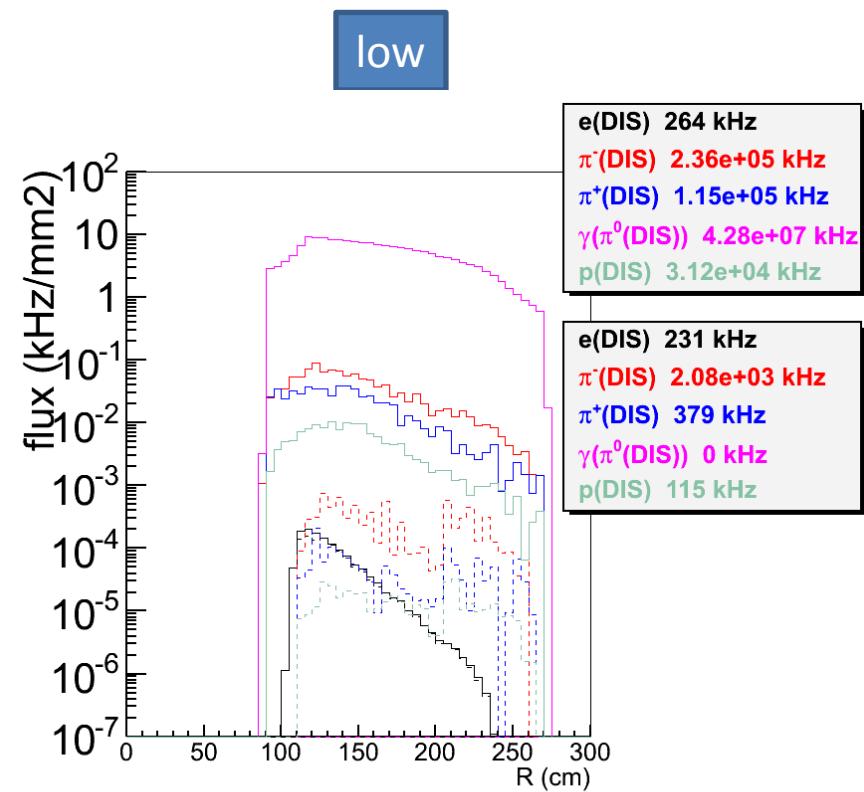
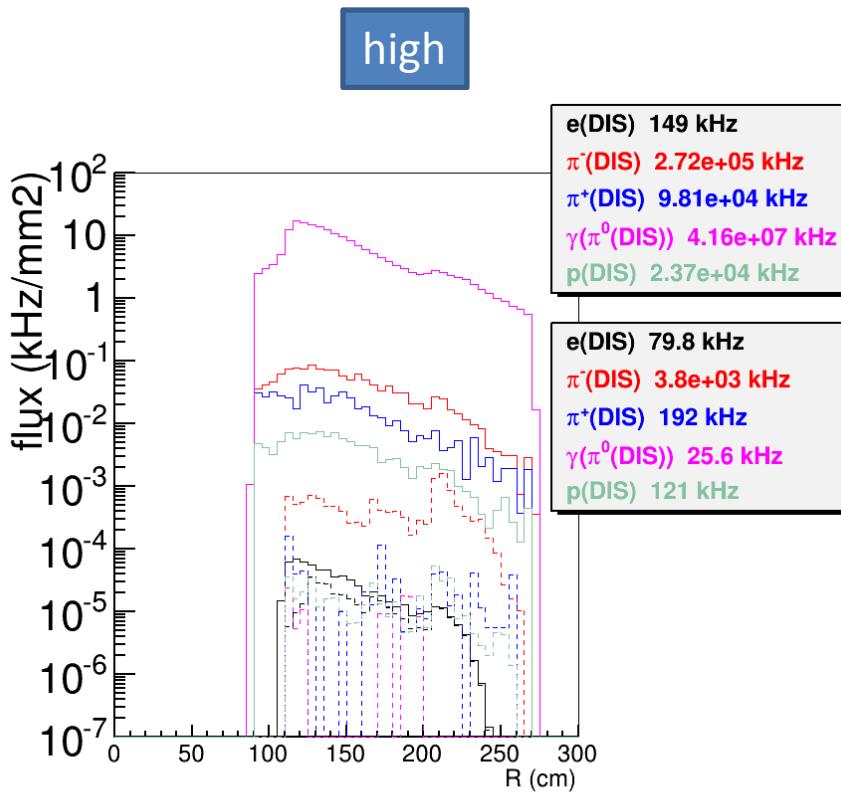


Rate before and after trig cut (More1 with more1 trig)



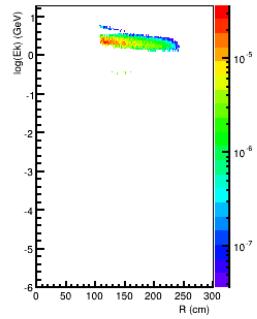
- Next slide shows the cut effect in “log(Ek) vs R”

Rate before and after trig cut (More1 block with more1 trig)

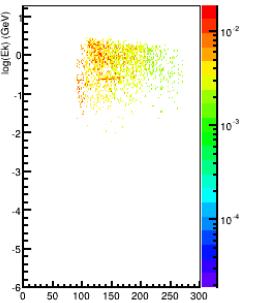


- Next slide shows the cut effect in “log(Ek) vs R”

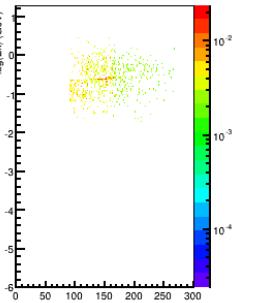
e DIS



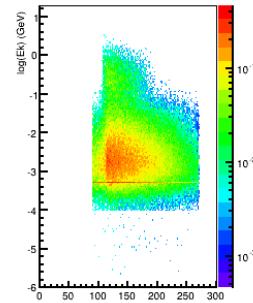
pi-



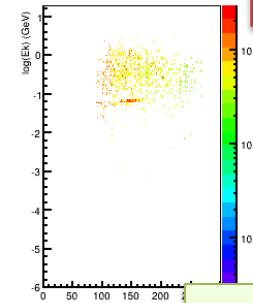
01+



$\gamma(\text{pi}0)$



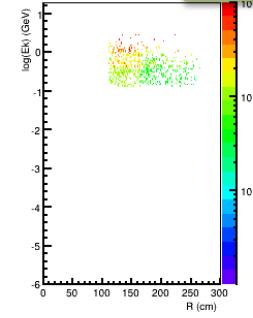
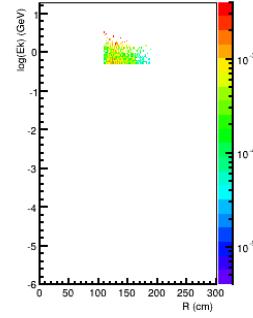
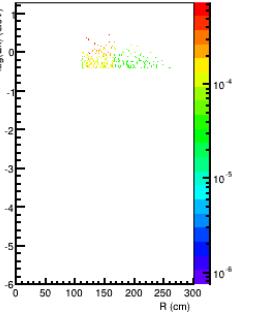
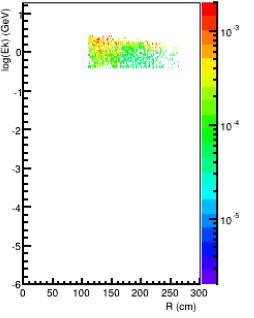
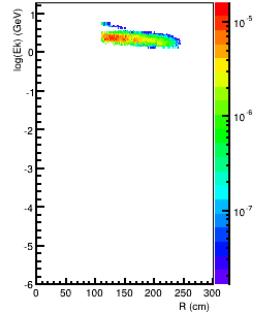
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More1
with
More1 trig

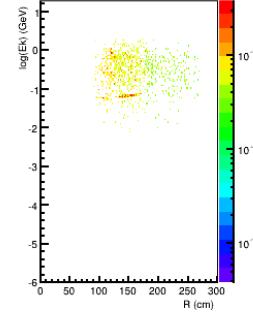
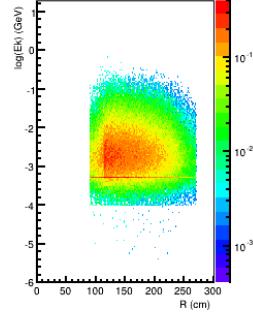
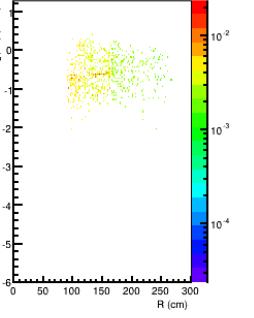
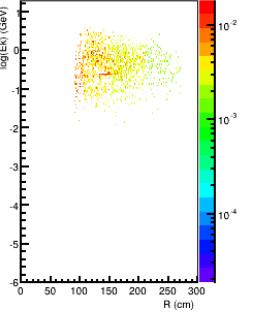
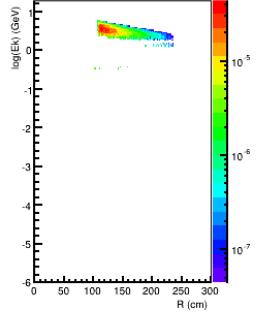
high,
Before
cut

Method 1

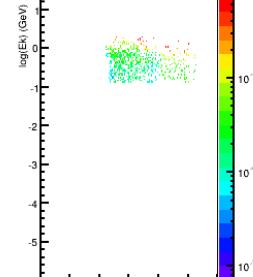
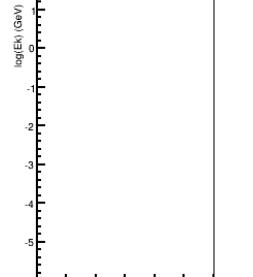
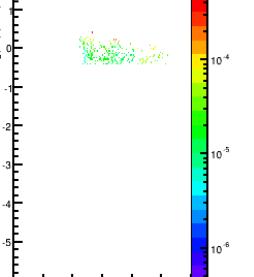
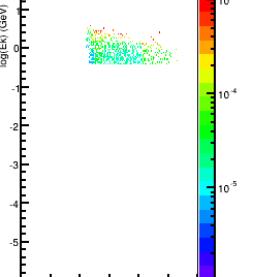
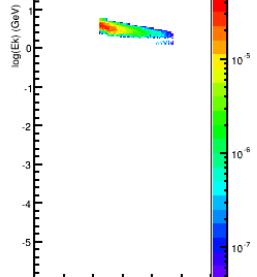


5
High,
After
cut

6

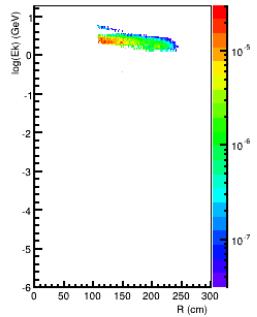


Low, Before cut

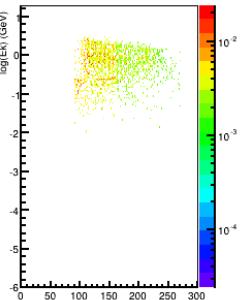


Low, After cut

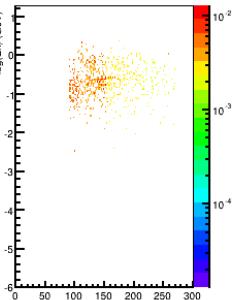
e DIS



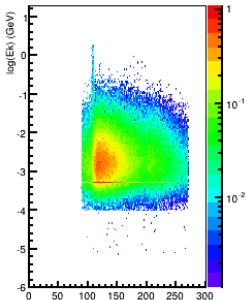
pi-



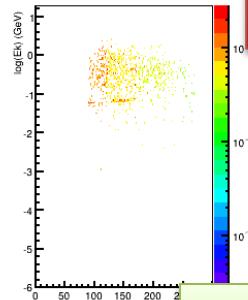
0i+



$\gamma(\text{pi}0)$

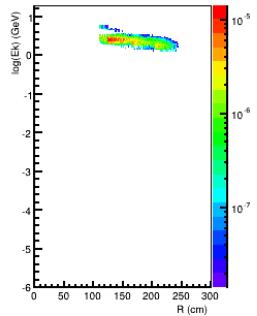


p



10^{-4}
high,
Before
cut

Method 1



High, After cut

A scatter plot showing the relationship between the logarithm of the energy E_k (in GeV) on the y-axis and the distance R (in cm) on the x-axis. The y-axis ranges from -6 to 7, and the x-axis ranges from 0 to 300. A color bar on the right indicates the density of points, ranging from 10^{-4} (dark blue) to 10^{-2} (red). The data points form a dense, roughly circular cluster centered around $R \approx 150$ cm and $\log(E_k) \approx 0$.

A scatter plot showing the relationship between the logarithm of the kinetic energy, $\log(E_k)$ (GeV), on the y-axis, and the radius, R (cm), on the x-axis. The x-axis ranges from 0 to 300 cm with major ticks every 50 units. The y-axis ranges from -6 to 0 GeV with major ticks every 1 unit. The data points are colored according to their density, with a color bar on the right ranging from blue (low density) to red (high density). The distribution shows a central concentration of high-density points (red) at $R \approx 150$ cm and $\log(E_k) \approx -1$ GeV, with a tail extending towards lower R values and higher $\log(E_k)$ values.

Low,
Before
cut

A scatter plot showing the relationship between the logarithm of the mass accretion rate ($\log(M^*/\text{Gyr})$) on the y-axis and the logarithm of the stellar mass ($\log(M^*/M_{\odot})$) on the x-axis. The data points are color-coded according to their density, with a color bar on the right ranging from 10^{-5} (blue) to 10^{-3} (red). The plot shows a dense concentration of points at higher stellar masses and higher mass accretion rates, with a more sparse distribution at lower values.

A scatter plot showing the relationship between $\log(E_t)$ (GeV) on the y-axis and $\log(S)$ on the x-axis. The x-axis ranges from approximately -1.5 to 1.5, and the y-axis ranges from -1.5 to 1.5. The data points are colored according to their density, with a color bar on the right ranging from 10^0 (blue) to 10^3 (red). The points form a roughly triangular cluster centered around $(\log(S) \approx 0.5, \log(E_t) \approx 0)$, with higher density (red) near the center and decreasing density (blue) towards the edges.

5
5
Low,
After
cut

Rate vs R

(more1 with more1 trig, lead)

Method 1

R(cm)	high					low				
	100-116	116-124	124-165	165-205	205-265	100-116	116-124	124-165	165-205	205-265
eDIS	1.38E+01	3.16E+01	8.90E+01	5.58E+01	1.53E+01	3.38E+01	7.18E+01	1.31E+02	2.35E+01	3.50E+00
π^-	1.24E+04	3.27E+04	1.42E+05	8.42E+04	4.47E+04	1.16E+04	2.42E+04	1.04E+05	6.17E+04	2.87E+04
π^+	5.26E+03	1.19E+04	5.15E+04	2.29E+04	1.26E+04	6.52E+03	1.19E+04	4.63E+04	3.04E+04	1.32E+04
$\gamma(\pi^0)$	1.08E+06	3.38E+06	1.29E+07	9.51E+06	6.91E+06	1.08E+06	3.14E+06	1.37E+07	1.31E+07	1.04E+07
p	7.53E+02	2.64E+03	1.22E+04	6.91E+03	3.92E+03	1.13E+03	2.62E+03	1.49E+04	6.03E+03	3.75E+03
total	1.10E+06	3.43E+06	1.31E+07	9.62E+06	6.97E+06	1.10E+06	3.18E+06	1.39E+07	1.32E+07	1.04E+07
eDIS	3.77E+00	1.39E+01	5.93E+01	3.64E+01	1.48E+01	2.45E+01	5.61E+01	1.24E+02	2.28E+01	3.34E+00
π^-	3.12E+02	1.43E+03	6.09E+03	1.91E+03	2.01E+03	1.52E+01	1.29E+02	7.73E+02	3.99E+02	6.13E+02
π^+	4.07E+01	1.55E+02	7.13E+02	1.66E+02	1.08E+02	2.47E+00	2.09E+01	1.08E+02	9.51E+01	1.30E+02
$\gamma(\pi^0)$	9.83E+02	3.14E+03	8.69E+03	2.97E+02	0.00E+00	0.00E+00	0.00E+00	0.00E+00	0.00E+00	0.00E+00
p	9.09E+00	4.24E+01	2.66E+02	7.99E+01	5.36E+01	3.54E-01	5.25E+00	4.27E+01	4.51E+01	6.15E+01
total	1.35E+03	4.78E+03	1.58E+04	2.49E+03	2.19E+03	4.25E+01	2.11E+02	1.05E+03	5.62E+02	8.08E+02

Rate vs R

(more 1 block with more 1 block trig, lead)

Method 1

R(cm)	high					low				
	100-116	116-124	124-165	165-205	205-265	100-116	116-124	124-165	165-205	205-265
eDIS	1.10E+01	2.49E+01	6.69E+01	3.04E+01	1.53E+01	3.39E+01	7.19E+01	1.32E+02	2.35E+01	3.49E+00
π^-	1.30E+04	2.90E+04	1.18E+05	6.73E+04	4.43E+04	9.58E+03	3.14E+04	1.02E+05	5.61E+04	3.66E+04
π^+	4.65E+03	1.13E+04	4.79E+04	2.06E+04	1.37E+04	5.11E+03	1.38E+04	5.61E+04	2.74E+04	1.27E+04
$\gamma(\pi^0)$	2.23E+06	6.26E+06	1.84E+07	7.83E+06	6.90E+06	1.20E+06	3.48E+06	1.44E+07	1.32E+07	1.04E+07
p	1.30E+03	2.46E+03	1.07E+04	5.59E+03	3.70E+03	1.30E+03	3.35E+03	1.58E+04	7.18E+03	3.62E+03
total	2.25E+06	6.30E+06	1.86E+07	7.92E+06	6.96E+06	1.22E+06	3.53E+06	1.46E+07	1.33E+07	1.05E+07
eDIS	1.01E+00	4.39E+00	3.75E+01	2.20E+01	1.49E+01	2.48E+01	5.55E+01	1.25E+02	2.29E+01	3.33E+00
π^-	1.21E+02	2.03E+02	7.77E+02	8.18E+02	1.89E+03	1.63E+01	1.16E+02	7.53E+02	3.52E+02	8.43E+02
π^+	2.87E+01	1.59E+01	1.02E+01	5.67E+01	8.07E+01	5.94E+00	4.96E+01	1.03E+02	5.38E+01	1.67E+02
$\gamma(\pi^0)$	4.36E+00	3.10E+00	0.00E+00	1.81E+01	0.00E+00	0.00E+00	0.00E+00	0.00E+00	0.00E+00	0.00E+00
p	6.38E+00	6.97E+00	2.58E+01	2.81E+01	5.34E+01	1.37E+00	3.83E+00	3.37E+01	2.45E+01	5.19E+01
total	1.61E+02	2.33E+02	8.51E+02	9.43E+02	2.04E+03	4.84E+01	2.25E+02	1.01E+03	4.53E+02	1.07E ₂₂ ⁰³

Rate (kHz)		More1 block(more1 trig)			More1 (CLEO trig)			More1 (more1 trig)		
		full	High	Low	full	High	Low	full	High	Low
e DIS	kry	413	148	265	469	205	264	same	same	same
π^-	Kry				5.29e4	4.11e4	1.19e4	same	same	same
	lead	5.08e5	2.72e5	2.36e5	5.45e5	3.15e5	2.30e5	same	same	same
	Kry				2.27e4	1.00e4	1.27e4	same	same	same
	lead	2.13e5	0.98e5	1.15e5	2.13e5	1.04e5	1.08e5	same	same	same
π^+	kry	5.06e3	5.06e3	0	1.20e6	1.20e6	0	same	same	same
	lead	8.44e7	4.16e7	4.28e7	7.51e7	3.37e7	4.14e7	same	same	same
	Kry				3.65e3	1.80e3	1.85e3	same	same	same
	lead	5.50e4	2.38e4	3.12e4	5.49e4	2.65e4	2.84e4	same	same	same
Total (lead)					Method 1					
e DIS	kry	335	104	231	403	145	258	359	128	231
π^-	Kry				5.22e3	4.00e3	1.22e3			
	lead	1.22e4	1.01e4	0.21e4	1.87e4	1.26e4	0.60e4	1.37e4	1.17e4	1.93e3
	Kry				0	0	0			
	lead	1.35e3	970	381	2.48e3	1.09e3	1.43e3	1.54e3	1.18e3	0.36e3
π^+	kry	2.3	2.3	0	3.22e4	3.22e4	0	1.23e4	1.23e4	0
	lead	230	115	0				1.31e4	1.31e4	0
	Kry				0	0	0			
	lead	927	404	120	957	475	481	0.61e3	0.45e3	0.16e3
Total (lead)		1.45e4	1.17e4	0.28e4	5.34e4	4.59e4	0.75e4	2.93e4	2.66e4	3.27e4